H0LiCOW: Measurement of H₀ from lensing

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Strong gravitationally lensed quasar

[Credit: ESA/Hubble, NASA]

Time delays

galaxy

distorted light rays

[Credit: ESA/Hubble, NASA]

Time delays

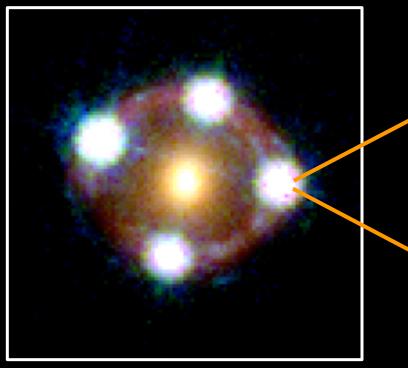


Map data @2016 GeoBasis-DE/BKG (@2009), Google 20 km L

[Credit: L. Huang, ASIAA]

Variability of quasar emission

HE0435-1223



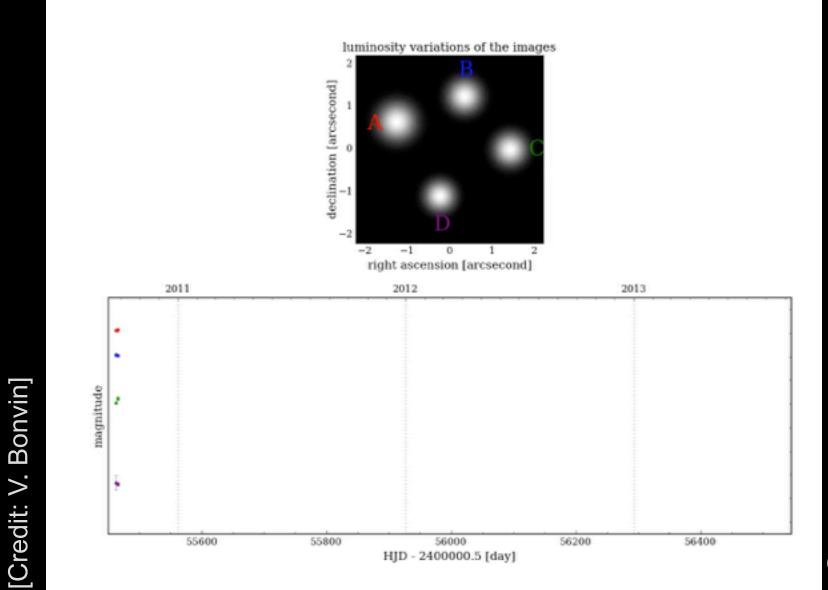
[Suyu et al. 2017]

quasar powered by accretion of material onto supermassive black hole:

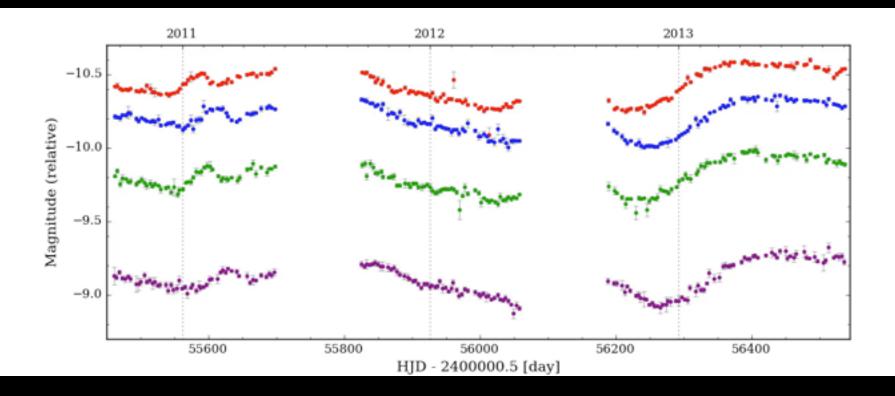


light emitted from quasar changes in time ("flickers")

Time delays

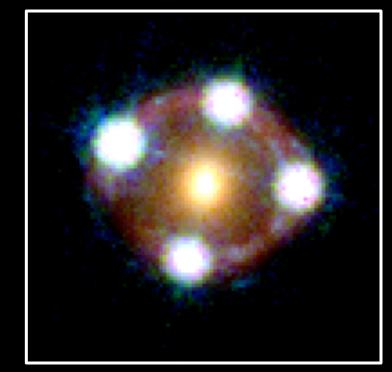


Time delays



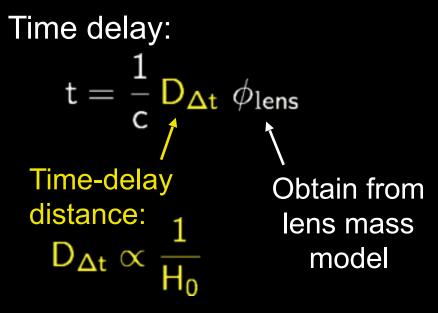
Cosmology with time delays

HE0435-1223



[Suyu et al. 2017]

Advantages:



For cosmography, need:

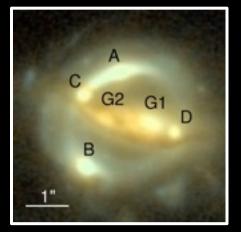
- (1) time delays
- (2) lens mass model
- (3) mass along line of sight
- simple geometry & well-tested physics

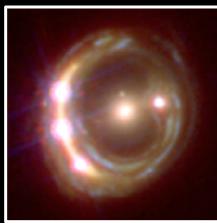
- one-step physical measurement of a cosmological distance

HOLICOW H₀ Lenses in COSMOGRAIL's Wellspring

B1608+656

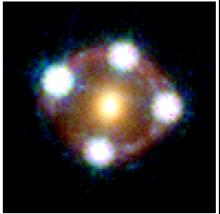
RXJ1131-1231



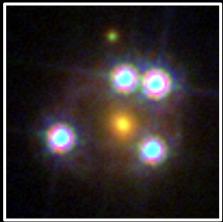


H₀ to <3.5% precision

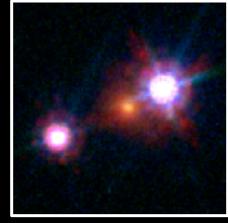
HE0435-1223



WFI2033-4723







[Suyu et al. 2017]

HOLiCOWers



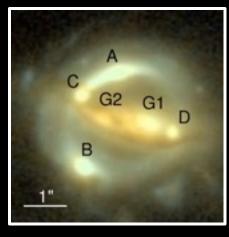


H0LiCOW: H₀ Lenses in COSMOGRAIL's Wellspring
→ Establish time-delay gravitational lenses as one of the best cosmological probes



B1608+656

[Suyu et al. 2010]



RXJ1131-1231



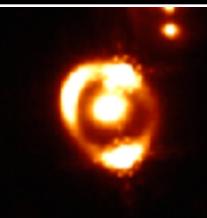
[Suyu et al. 2013, 2014; Tewes et al. 2013]

HE0435-1223

[Wong, Suyu et al. 2017; Rusu et al. 2017; Sluse et al. 2017; Bonvin et al. 2017]



SDSS1206+4332

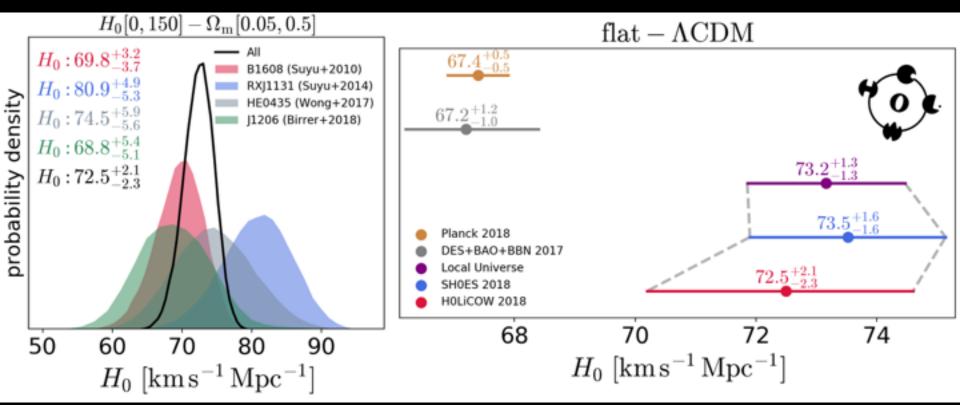


part of extended sample

[Birrer et al. 2018]

H₀ from 4 strong lenses

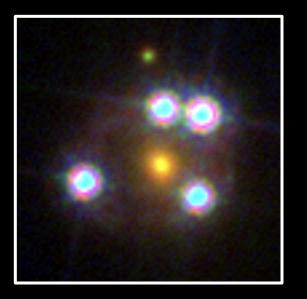
Blind analysis to avoid confirmation bias



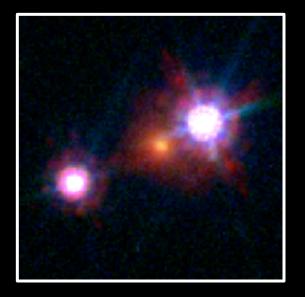
H_0 with 3% precision in flat Λ CDM

Looking forward

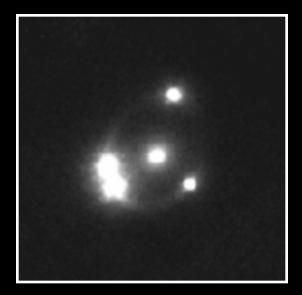
WFI2033-4723



HE1104-1805



WFI2033-4723: blind analysis ongoing [Rusu et al., Sluse et al., Wong et al., in prep.]



PG1115+080

PG1115+080 : blind analysis of adaptive-optics and HST images ongoing [Chen et al., in prep.]³

Towards hundreds of lenses

Hyper Suprime-Cam Survey



8m Subaru Telescope Mauna Kea, Hawaii

- 1400 deg² with i_{limit}~26
- 2014-2019
- expect ~600 lenses
 [Oguri & Marshall 2010]

Dark Energy Survey

STRong-lensing Insights into Dark Energy Survey (PI: Treu) 4m Blanco Telescope, CTIO, Chile

- 5000 deg² with i_{limit}~24
 2012-2017
- expect ~1100 lenses
 [Oguri & Marshall 2010]

Kilo Degree Survey

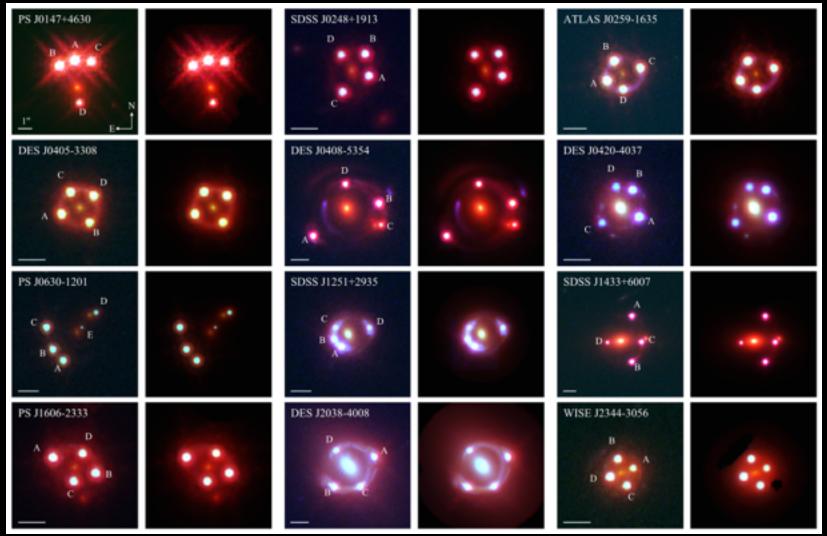


2.6m VLT Survey Telescope, Paranal, Chile

1500 deg² with r_{limit}~25
2011-~2018

New quads imaged with HST

New lens systems discovered in DES, Pan-STARRS, SDSS, ATLAS:



[Shajib et al. 2018]

Strongly lensed supernova

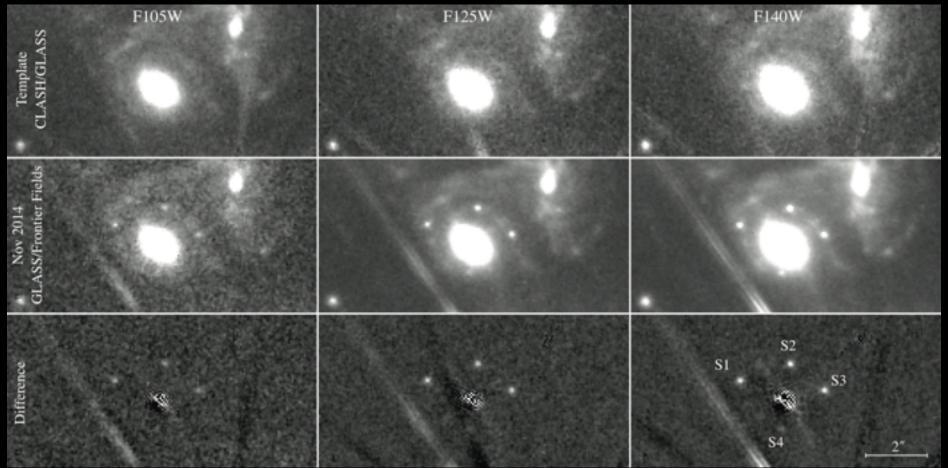


MACS 1149.6+2223

[Kelly et al. 2015] ¹⁶

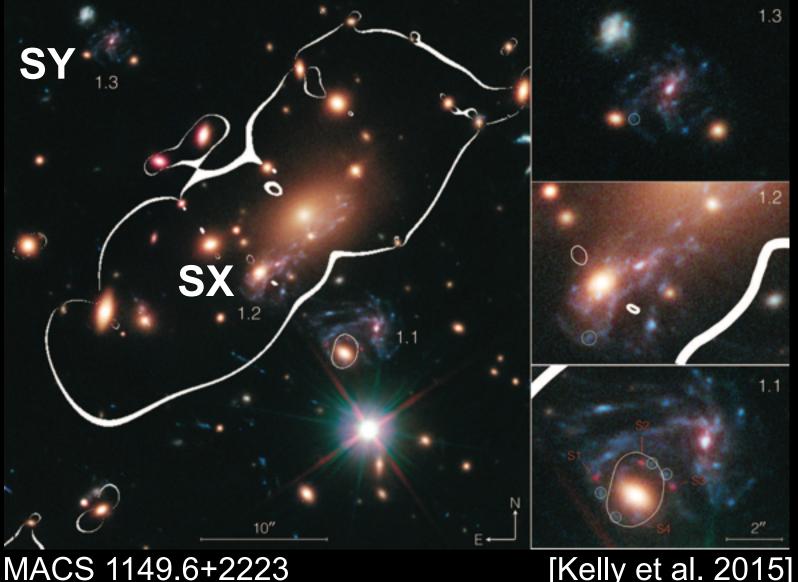
Supernova "Refsdal"

discovered serendipitously in November 2014



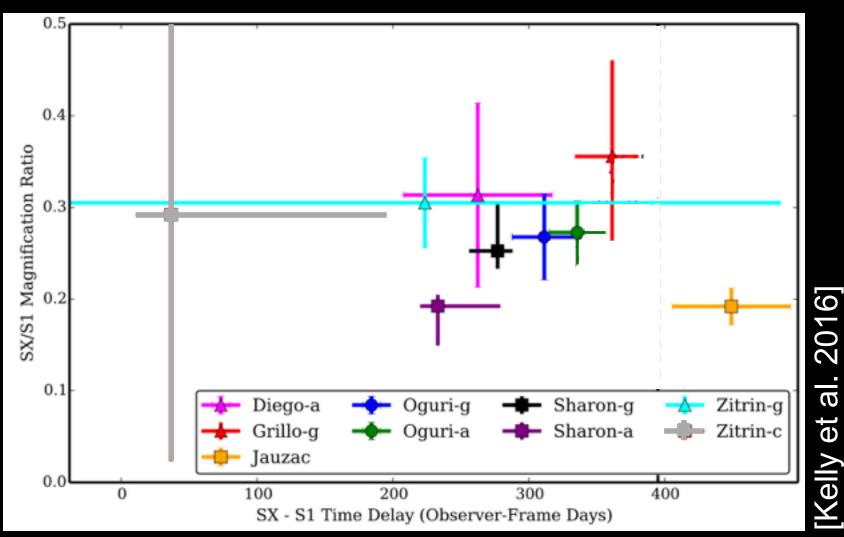
[Kelly et al. 2015] ₁₇

When will the other SN images appear?

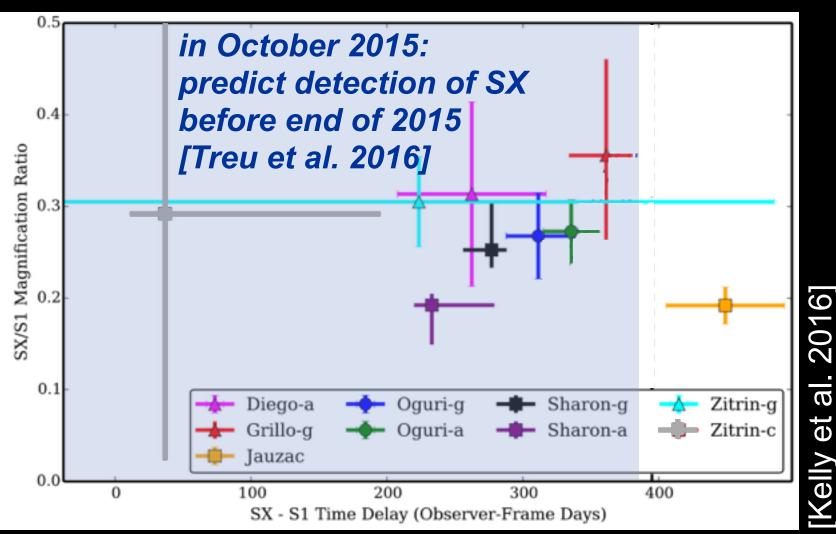


[Kelly et al. 2015] ¹⁸

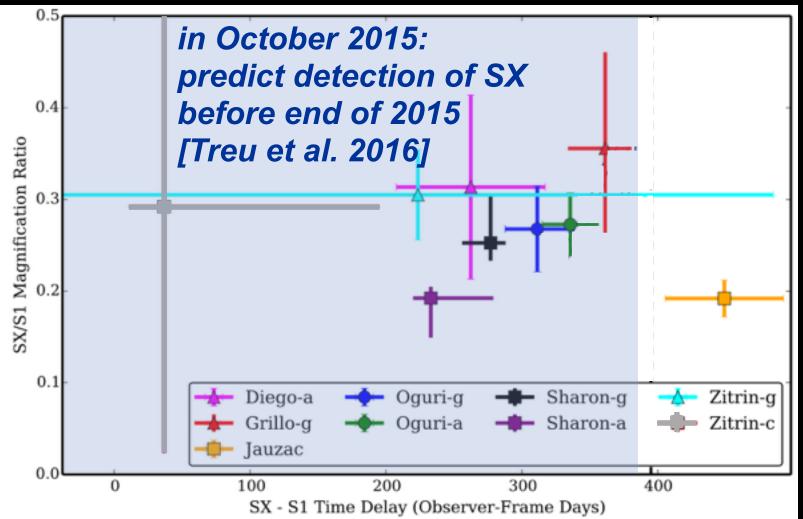
Predicted magnification and delay



Predicted magnification and delay

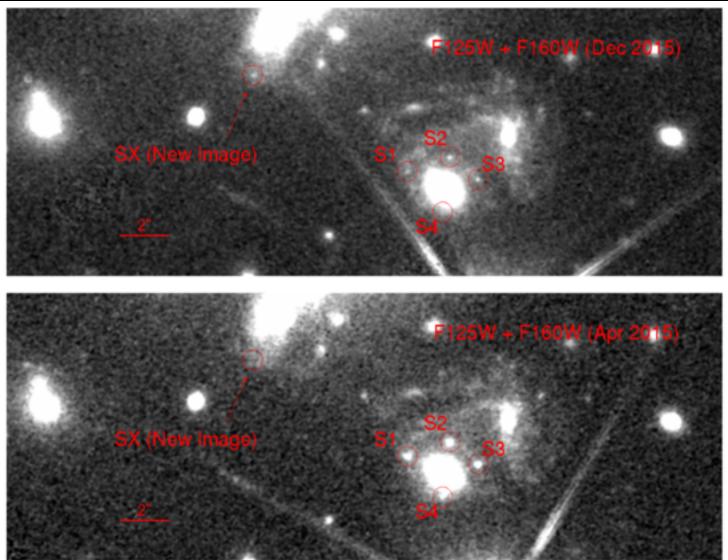


Predicted magnification and delay

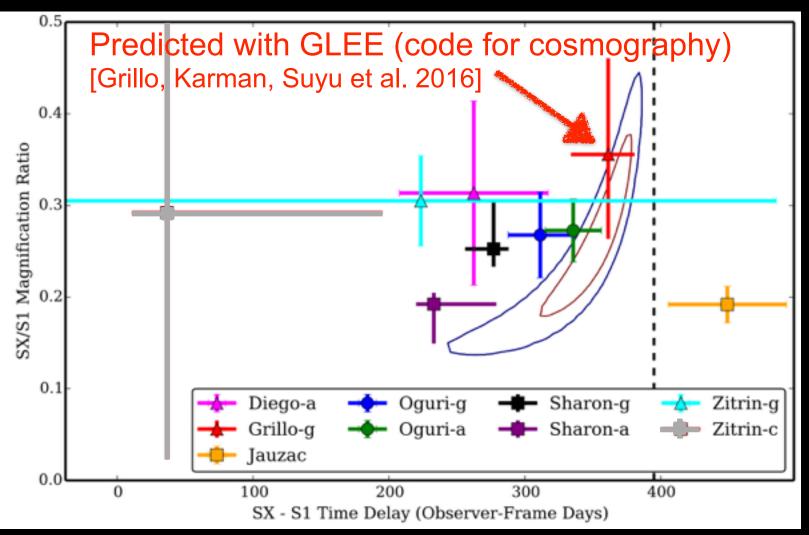


HST observations in Oct 2015: no sign of SX in Nov 2015: no sign of SX...

Appearance of image SXDecember 2015[Kelly et al. 2016]

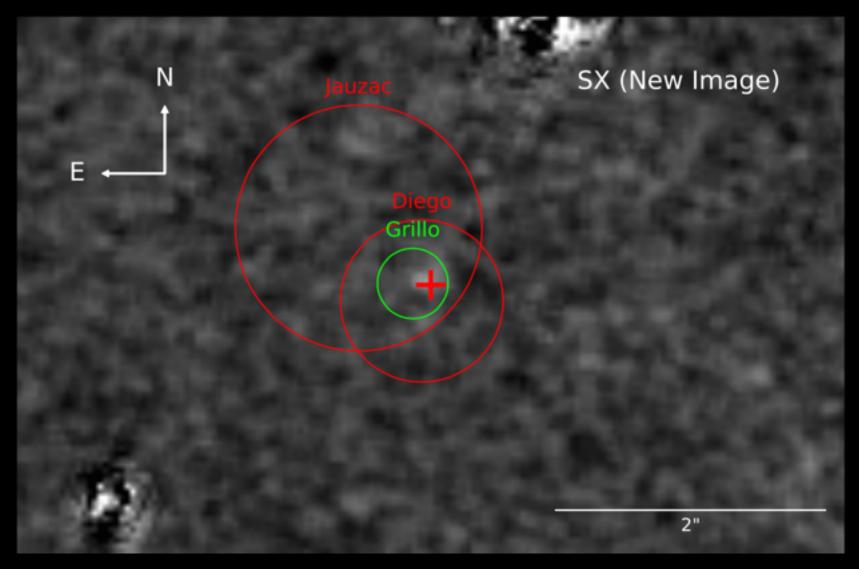


Magnification and delay



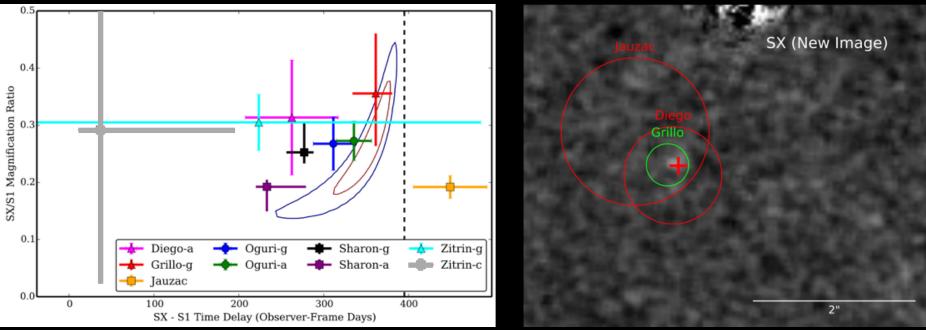
[Kelly et al. 2016] 23

Spot on!



[Kelly et al. 2016] ²⁴

Successful prediction

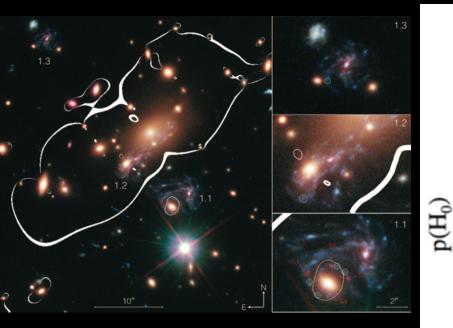


[Kelly et al. 2016]

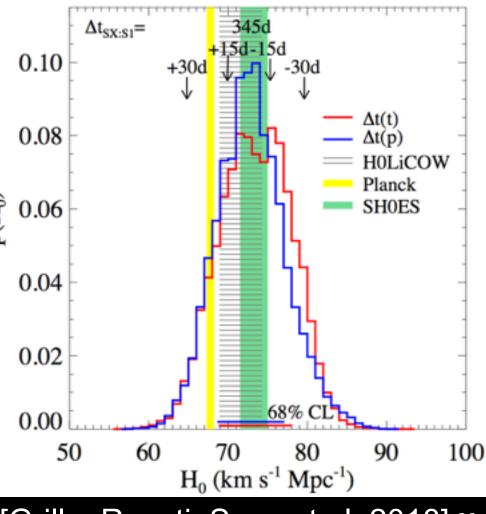
- Grillo et al. predicted successfully and precisely all three observables of image SX: time, magnification and location
- GLEE [Suyu & Halkola 2010] accurately and precisely predicted the appearance of SX as a result of its ability to reconstruct the entire SN host galaxy

H₀ à la Supernova Resfdal

feasibility study of using SN Refsdal for H₀ measurement



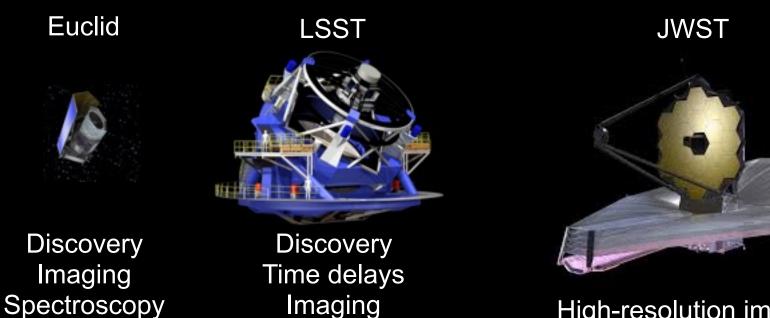
- S1-S2-S3-S4 delays from Rodney et al. (2016)
- SX-S1 delay estimated based on detection in Kelly et al. (2016)



[Grillo, Rosati, Suyu et al. 2018] ²⁶

Future Prospects

Experiments and surveys in the 2020s including Euclid and Large Synoptic Survey Telescope (LSST) will provide ~10,000 lensed quasars and ~100 lensed supernovae [Oguri & Marshall 2010]



High-resolution imaging & spectroscopy

Summary

- Time-delay distances $D_{\Delta t}$ of each lens can be measured with uncertainties of ~5-8% including systematics
- From 4 lenses in H0LiCOW, $H_0 = 72.5^{+2.1}_{-2.3}$ km/s/Mpc in flat ACDM, a 3% precision measurement independent of other probes
- Search is underway to find new lenses in DES, HSC, PanSTARRS, and Gaia surveys
- SN Refsdal blind test demonstrated the robustness of our lens mass modeling approach
- Current and future surveys will have thousands of new time-delay lenses, providing an independent and competitive probe of cosmology

Thank you!